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#### Validation of Himawari-9/AHI Level-1 and -2 data during In-orbit Test

The Himawari-9 geostationary meteorological satellite of the Japan Meteorological Agency (JMA) was launched on 2 November 2016 and put into in-orbit standby as backup for Himawari-8 on 10 March 2017. Himawari-9 features the Advanced Himawari Imager (AHI), which is identical to the AHI on board Himawari-8. As accurate image navigation and radiometric calibration are essential in leveraging the imager's potential, Himawari-9/AHI Level-1 and -2 data performance was validated during the satellite's period of in-orbit test (IOT).

Image navigation and registration errors determined from observation during the IOT phase were in the same order as those of Himawari-8/AHI. The validation results for Himawari-9/AHI calibration generally show close correspondence to those for Himawari-8/AHI, although larger biases are seen in several bands. Further evaluation will be performed in the near future.

Quality assessment of Atmospheric Motion Vectors (AMVs) and Clear Sky Radiance (CSR) for the AHIs on board Himawari-8 and -9 was also carried out. The rawinsonde statistics for AMVs were similar for both satellites. In regard to CSR, the first-guess (FG) departure statistics for B08 and B10 were similar for both satellites, although the B09 FG departure for Himawari-9 was 0.2 K greater.

Action/Recommendation proposed: none

#### Validation of Himawari-9/AHI Level-1 and -2 data during its In-Orbit Test

#### 1 INTRODUCTION

The Japan Meteorological Agency's new-generation Himawari-8 geostationary meteorological satellite began operation in July 2015 after in-orbit test (IOT) and checking of the overall system. The identically configured Himawari-9 was launched on 2 November 2016, and was put into in-orbit standby as backup for Himawari-8 on 10 March 2017 after IOT. This set-up will help to ensure the stability of satellite observation for the East Asia and Western Pacific regions for 15 years. Himawari-8 and -9 feature new Advanced Himawari Imager units (referred to here as AHI-8 and AHI-9) with a sensor configuration similar to that of the Advanced Baseline Imager (ABI) on the GOES-16 satellite (Table 1-1).

The AHI conducts observation to produce full-disk imagery every 10 minutes. During the 10-minute timeline of such observation, the AHI scans the Japan and Target Area observations every 2.5 minutes and two Landmark Area observations at 30-second intervals. One of the Landmark Area observations can be applied to the moon based on advance prediction of lunar positioning. The large amount of AHI lunar observation data produced (in comparison to that of heritage imagers on board geostationary satellites) is expected to be useful in validating AHI calibration trends.

Section 2 of this paper provides general information on the Himawari-8/9 AHI, Section 3 covers the image navigation and registration (INR) performance of Himawari-9, Section 4 details the results of validation regarding radiometric calibration, Section 5 outlines quality assessment of AMVs and CSRs for Himawari-9, and Section 6 provides the conclusion.

	Himawari-8/-9	GOES-16	MTSAT-2
Radiometer	AHI	ABI	IMAGER
Bands	16	16	5
Banus	VIS:3, NIR:3, IR:10	VIS: 2, NIR: 4, IR:10	VIS:1, IR:4
Spatial resolution at SSP	0.5 km, 1 km, 2 km	0.5 km, 1 km, 2 km	1km, 4 km
Full disk observation	Every 10 minutes	Every 15 minutes (planned)	Hourly
Scan pattern	West to East 23 scans cover full disk	West to East 22 scans cover full disk	Both directions from West and East. Around 1400 scans cover full disk
Orbit determination	Ranging	Ranging + GPS	Ranging
Attitude determination	Star tracker + Rate	Star tracker + Rate	Sun sensor + Earth
	sensor and so forth	sensor and so forth	sensor
Navigation correction	Landmark analysis	Star sensing	Landmark analysis

Table 1-1 General information on Himawari-8/-9 AHI, GOES-16 ABI and MTSAT-2 IMAGER

#### 2 AHI BASICS

Table 2-1 and Figure 2-1 show observing band configuration and the spectral response functions (SRFs) of AHI-8 and -9, respectively. The AHIs are manufactured by the ABI vendor, but feature the 0.51- $\mu$ m band rather than the 1.38- $\mu$ m band. True-color imagery is derived by composing the three visible bands (i.e., blue: 0.47  $\mu$ m; green: 0.51  $\mu$ m; red: 0.64  $\mu$ m). SRF data are provided on JMA's Meteorological Satellite Center (MSC) website at:

http://www.data.jma.go.jp/mscweb/en/himawari89/space\_segment/spsg\_ahi.html

		B01	B02	B03	B04	B05	B06	B07	B08	B09	B10	B11	B12	B13	B14	B15	B16
	Wlen	0.47	0.51	0.64	0.86	1.6	2.3	3.9	6.2	6.9	7.3	8.6	9.6	10.4	11.2	12.4	13.3
	Sres	1	1	0.5	1	2	2	2	2	2	2	2	2	2	2	2	2
MTSAT-2	Wlen			0.68				3.7	6.8					10.8		12.0	
	Sres			1				4	4					4		4	

Table 2-1 Band configuration of AHI-8/-9 and IMAGER on MTSAT-2. When is the central wavelength  $[\mu m]$ , and Sres represents spatial resolution at the SSP [km].

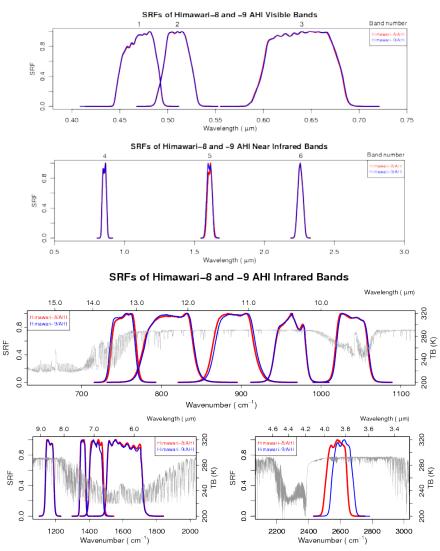


Figure 2-1 Spectral response functions of AHI-8 (red) and -9 (blue) Page 3 of 13

#### 3 IMAGE NAVIGATION AND REGISTRATION

The INR process for AHI-9 is identical to that of AHI-8, which is based on information from star trackers, inertia reference units and angular rate sensors on board the satellite. JMA validated the INR performance of AHI-9 during in-orbit test using the landmark and band-to-band co-registration analysis methods. The former involves comparing coastlines recognized in observation data with those on reference maps to derive absolute INR errors. The latter is similar, but involves comparison of cloud patterns as well as coastlines in target-band observation data with those in the reference band (e.g., B13) for estimation of band-to-band relative INR errors in each band with the reference band.

Figure 3-1 shows landmark analysis results for B13 Himawari Standard Data (i.e., Level-1 data after rectification). The estimated magnitudes of image navigation errors are around 7.37  $\mu$ rad (0.26 km at the sub-satellite point (SSP)) and 10.5  $\mu$ rad (0.38 km at the SSP) at 00:00 and 12:00 UTC, respectively.

Table 3-1 shows diurnally averaged INR errors based on landmark analysis for several bands on the same day. The results indicate that AHI-9 image navigation performance is comparable to that of AHI-8. However, several significant navigation errors are observed daily in AHI-9 observation imagery (figures not shown). The root cause of this is currently being researched by the satellite vendor.

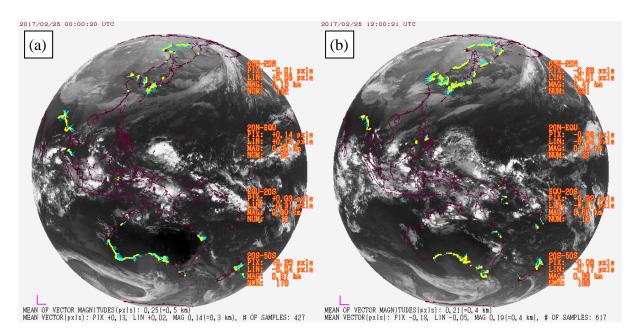


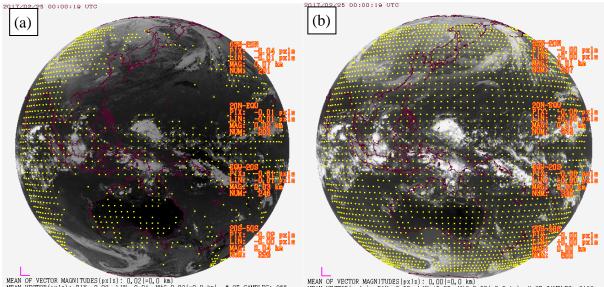
Figure 3-1 (a) AHI-9 Landmark analysis results for B13 at 00:00 on UTC 25 Feb. 2017 (b) As per (a), but for 12:00 UTC. Yellow circles show reference points used to compute misplacement. Blue segments connected with the yellow circles indicate directions and magnitudes of misplacement at each reference point. The short magenta segment (bottom-left) represents the unit length of magnitude for misplacement (i.e., 56 µrad, or one pixel for infrared bands, for which the horizontal resolution is 2 km at the SSP).

	B01	B07	B13	B15
East-West ( $\mu$ Rad)	0.026	-4.145	-3.309	-3.329
North-South ( $\mu$ Rad)	-3.461	-0.241	1.287	3.134
Bias Magnitude( $\mu$ Rad)	9.100	10.313	9.633	10.531

Table 3-1 Diurnally averaged INR errors based on landmark analysis for several AHI-9 bands on 25 Feb 2017

Band-to-band co-registration correction has been applied since the start of AHI-9 inorbit test. Figure 3-2 shows the co-registration analysis results for B07 and B15 with respect to B13. The magnitude of the misplacement is below 1.1 µrad (0.04 km at the SSP), and the co-registration performance of other bands is comparable.

Table 3-2 shows diurnally averaged band-to-band co-registration errors for all bands based on observation data for 25 February 2017. As with image navigation performance, AHI-9 co-registration performance is comparable to that of AHI-8.



MEAN OF VECTOR MAGNITUDES(pxls): 0.02(=0.0 km) MEAN VECTOR(pxls): PIX -0.02, LIN -0.01, MAG 0.02(=0.0 km), ♯ OF SAMPLES: 966

MEAN OF VECTOR MAGNITUDES(pxls): 0,00(=0,0 km) MEAN VECTOR(pxls): PlX -0,00, LIN +0,00, MAG 0,00(=0,0 km), ♯ OF SAMPLES: 2162

Figure 3-2 (a) band-to-band co-registration analysis for B07 in comparison with B13 at 00:00 UTC on 25 Feb. 2017, and (b) the same for B15. Circles and segments show reference points and their magnitudes as in Figure 3-1.

Table 3-2 Diurnally averaged bar	nd-to-band co-registration errors for all bands with
respect to B13 for 25 Februar	y 2017

Band number	B01	B02	B03	B04	B05
Bias Magnitude( $\mu$ Rad)	0.990	1.014	1.220	1.553	1.313
Band number	B06	B07	B08	B09	B10
Bias Magnitude( $\mu$ Rad)	1.327	0.817	0.289	0.444	0.501
Band number	B11	B12	B14	B15	B16
Bias Magnitude( $\mu$ Rad)	0.576	0.519	0.579	0.223	0.286

#### 4 CALIBRATION

For calibration of observation data, the AHI has a solar diffuser serving as a solar calibration target for visible and near-infrared bands (i.e., B01-06) and a blackbody serving as an internal calibration target for 10 infrared bands (i.e., B07-16). JMA has been validating AHI-8 data quality based on the GEO-LEO technique (involving intercalibration and vicarious calibration (Okuyama et al. 2015)), lunar calibration and other approaches. These calibration and validation methods have been developed via international collaboration with NOAA (Wu 2016), EUMETSAT and GSICS member agencies in addition to collaborative research with the Atmosphere and Ocean Research Institute at the University of Tokyo. This article reports on AHI-9 validation using some of these approaches. Tables 4-1 and 4-2 summarize the results for all 16 AHI-8 and -9 bands.

For visible and near-infrared bands (Table 4-1), radiances were validated based on 1) comparison with top-of-atmosphere radiance computed via radiative transfer simulation (vicarious calibration) and 2) a ray-matching approach with reference to S-NPP/VIIRS. Estimated radiance biases of AHI-9 from the vicarious calibration approach were +2.94 and -5.48% for B01 and B06, but the biases for other bands were less than +/- 2.0%. The ray-matching approach provided results consistent with those of the vicarious calibration approach for B01 and B06. Infrared inter-calibration (Table 4-2) with reference to hyperspectral infrared sounders such as Metop-A/IASI showed that brightness temperature biases for AHI-9 are in the same order as those validated for AHI-8 (less than 0.25 K for standard scenes (i.e., simulated brightness temperature for the US standard atmosphere)) in all 10 infrared bands.

The frequent full-disk observation conducted by AHI-8 and -9 (with a repeat cycle of 10-minutes) also enables application of the highly useful GEO-GEO comparison approach. Although the GEO-GEO approach is a relative comparison method without accurate reference sensor such as IASI and VIIRS, the huge amounts of collocated data enable identification of calibration issues (such as diurnal variation of biases, stray light and banding) on a real-time basis. In this study, AHI-8 and -9 Himawari Standard Data from the same observation time and the same band were averaged for areas of 19 x 19 pixels and compared in terms of scaled radiance (for visible and near-infrared bands) and brightness temperature (for infrared bands).

For verification of consistency between the GEO-LEO and GEO-GEO approaches mentioned above, the relative differences between AHI-8 and -9 of the GEO-LEO approaches are shown at the bottom of Tables 4-1 and 4-2 along with GEO-GEO approach results. Ratios of AHI-9/AHI-8 are shown for visible and near-infrared bands in Table 4-1, and differences in biases for AHI-9 minus AHI-8 in standard scenes for infrared bands are shown in Table 4-2. For infrared bands, the effects of spectral response function difference were removed from GEO-GEO results using AHI-8 and -9 pseudo data from Metop-A/IASI for a particular scene (14 January 2015, latitude within 30 deg., longitude within 80 deg. from the Himawari-8/-9 subsatellite point (140.7°E, 0.0°N)). Close correspondence between the comparison results for each approach is observed in Figure 4-1. This validation approach offers a promising solution for the generation of a new inter-calibration product combining GEO-LEO and GEO-GEO comparison results to account for diurnal calibration variations.

Table 4-1 Estimated observation biases of AHI-8 and -9 for visible and near-infrared bands (in scaled radiance [%] for vicarious calibration and GEO-GEO approaches, and in reflectance [%] for ray-matching approach). Observation data for the period from 14 to 28 February 2017 for AHI-8 and AHI-9 are used in these approaches. AHI-9 calibration coefficients other than offset terms were determined based on pre-launch ground test whereas AHI-8 calibration slopes were updated on 8 June 2015 to reflect the solar diffuser viewing data collected in orbit.

		B01	B02	B03	B04	B05	B06
AHI-8 [%]	Vicarious cal.	-1.61	-2.55	+0.71	+1.62	+4.61	-3.02
Апі-о [%]	Ray-matching	-2.50	+1.13	+1.92	+0.62	+6.59	-4.50
	Vicarious cal.	+2.94	-1.85	-1.82	+0.01	-1.15	-5.48
AHI-9 [%]	Ray-matching	+3.15	+2.83	+0.39	-0.031	+0.59	-6.22
	Vicarious cal.	+4.63	+0.72	-2.51	-1.58	-5.51	-2.54
(AHI-9)/(AHI-8) [%]	Ray-matching	+5.79	+1.68	-1.50	-0.65	-5.63	-1.80
	GEO-GEO	+4.75	+1.02	-2.22	-1.45	-5.79	-2.00

Table 4-2 Estimated observation biases of AHI-8 and -9 in brightness temperature [K] for infrared bands. Observation data for the period from 14 to 28 February 2017 for AHI-8 and AHI-9 are used in these approaches. Only results from nighttime Metop-A/IASI (10 - 14 UTC) are shown.

-		B07	B08	B09	B10	B11	B12	B13	B14	B15	B16
AHI-8 [K]	Metop-A/IASI	-0.13	-0.24	-0.25	-0.14	-0.09	-0.25	-0.01	0.00	-0.08	+0.05
AHI-9 [K]	Metop-A/IASI	-0.08	-0.25	-0.04	-0.15	-0.12	-0.20	-0.10	-0.11	-0.13	-0.23
(AHI-9)-	Metop-A/IASI	+0.05	-0.01	+0.21	-0.01	-0.03	+0.05	-0.09	-0.11	-0.04	-0.28
(AHI-8)[K]	GEO-GEO	+0.05	-0.03	+0.19	-0.02	-0.02	+0.06	-0.05	-0.08	-0.02	-0.25

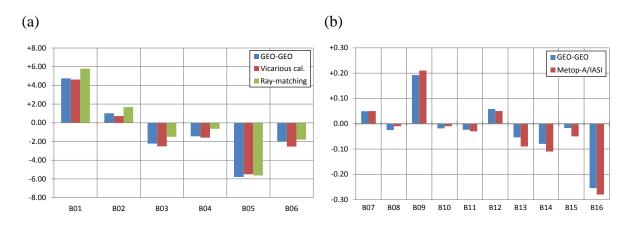


Figure 4-1 Consistency of different approaches in terms of (a) ratios of (AHI-9)/(AHI-8) [%] for visible and near-infrared bands, and (b) differences in AHI-9 and AHI-8 biases [K] for infrared bands

#### 5 LEVEL-2 PRODUCTS

Quality assessment of AMVs (Atmospheric Motion Vectors) and CSR (Clear Sky Radiance) in the Himawari-9 IOT phase showed consistency with products from the Himawari-8 IOT phase.

#### 5.1 **AMV**

Rawinsonde statistics for Himawari-8 and -9 AMVs were assessed for the period from 1st to 26th February 2017. The Mean Vector Difference (MVD), RMS Vector Difference (RMSVD) and wind speed bias (BIAS) exhibited improvement, but variations between Himawari-8 and -9 were small. The method, reporting period and filters are detailed below.

Rawinsonde comparison statistics Method: Comparison of rawinsonde wind data with AMV winds within a 150-km radius of a RAOB site Reporting period: 1st – 26th February 2017 Filters: VERT. DIST. (> = 700 hPa) < 50 (hPa) VERT. DIST. (> = 700 hPa) < 35 (hPa) QUALITY >= 85 0.5 \* 0.5 deg. latitude/longitude grid point data SPEED DIFF. < 30 (m/s) – DIRECTION DIFF. < 90 (deg)

The statistics for B03, B08, B09, B10 and B13 are shown in Tables 5-1 to 5-5. Red and blue represent improvement and degradation, respectively.

Table 5-1 VIS (B03) rawinsonde statistics for Himawari-8 (a) and Himawari-9 (b)

(a)					(b)				
. ,	Sele	llite in use : cted channel : et Type :		Himawari-8 B03 (0.64 um) cloudy		Selec	llite in use : sted channel : et Type :		Himawari-9 B03 (0.64 um) cloudy
	ALL REGIONS	NH EX-TROP	TROP	SH EX-TROP		ALL REGIONS	NH EX-TROP	TROP	SH EX-TROP
BIAS SPD NCMV NC HIGH LEVEL	5.07 -0.37 15.80 388324 21117	5.50 -0.49 20.98 102042 8013	4.76 -0.32 12.26 166196 11014	4.09 4.90 -0.19 14.60 120086 2090	BIAS SPD NCMV NC HIGH LEVEL	19907	-0.44 21.49 81639 8517	4.70 -0.28 12.21 135165 9506	4.63 -0.18 15.23 93343 1884
MVD RMSVD BIAS SPD NCMV NC	5.64 -0.83 20.46 81775	6.75 -0.75 37.00 13123	4.34 5.12 -0.81 14.12 51568 6922	5.31 6.17 -1.32 25.09 17084 686	MVD RMSVD BIAS SPD NCMV NC	4.66 5.59 -0.63 21.24 69486 9494	5.42 6.57 -0.38 36.36 11702 2734	4.27 5.05 -0.69 13.99 42998 6090	5.14 5.93 -1.13 25.40 14786 670
MEDIUM LEVEL MVD RMSVD BIAS SPD NCMV NC	4.15 4.97 -0.55 16.45 54138	5.39 -0.49 21.65 19798		4.35 5.02 -0.64 13.50 17622 212		4.18	4.49 5.32 -0.50 20.88 15586 2312	3.64 4.24 -0.46 8.22 13384 1071	3.67 4.31 -0.38 12.74 13456 272
LOW LEVEL MVD RMSVD BIAS SPD NCMV NC	3.53 4.29 0.28 9.62 252411 7932	3.71 4.54 -0.32 10.03 69121 3717	3.37 4.09 0.91 9.46 97910 3023	3.35 3.95 0.54 8.76 85380 1192	BIAS SPD	3.29 4.02 0.15 9.70 198235 6758	3.47 4.23 -0.45 10.18 54351 3471	3.89 0.86 9.39 78783	3.00 3.53 0.55 8.72 65101 942

(a)	•	,		(	(b)					
	Selec	llite in use : sted channel : et Type :		Himawari-8 B08 (6.2 um) cloudy	( )	Sele	llite in use : cted channel : et Type :		Himawari-9 B08 (6.2 um) cloudy	
	ALL REGIONS	NH EX-TROP	TROP	SH EX-TROP		ALL REGIONS	NH EX-TROP	TROP	SH EX-TROP	
ALL LEVEL					ALL LEVEL					
MVD				5.45		4.71		4.42		
				6.40	RMSVD	5.60	6.51	5.22	5.88	
BIAS	-0.12	-0.12	-0.13	-0.01		-0.04		-0.09	0.01	
SPD	21.99	37.40	17.20	28.33	SPD	22.43	36.52	17.03	28.24	
NCMV	523670	69019	358394	96257	NCMV	399883	54134	274609	71140	
NC	62355	13466	46498	2391	NC	53319	13599	37689	2031	
HIGH LEVEL					HIGH LEVEL					
MVD	4.77	5.64	4.48	5.45	MVD	4.71	5.48	4.42	5.04	
RMSVD	5.69	6.81	5.28	6.40	RMSVD	5.60	6.51	5.22	5.88	
BIAS	-0.12	-0.12	-0.13	-0.01	BIAS	-0.04	0.11	-0.09	0.01	
SPD	21.99	37.40	17.20	28.33	SPD	22.43	36.52	17.03	28.24	
NCMV	523670	69019	358394	96257	NCMV	399883	54134	274609	71140	
NC	62355	13466	46498	2391	NC	53319	13599	37689	2031	
MEDIUM LEVEL					MEDIUM LEVE	r.				
MVD	0.00		0.00	0.00	MVD	0.00	0.00	0.00	0.00	
RMSVD	0.00	0.00	0.00	0.00	RMSVD	0.00	0.00	0.00	0.00	
BIAS	0.00		0.00	0.00	BIAS	0.00	0.00	0.00	0.00	
SPD	0.00	0.00	0.00	0.00	SPD	0.00	0.00	0.00	0.00	
NCMV	0	0	0	0	NCMV	0	0	0	0	
NC	0	0	0	0	NC	Ō	Ō	0	0	
LOW LEVEL					LOW LEVEL					
MVD	0.00	0.00	0.00	0.00	MVD	0.00	0.00	0.00	0.00	
RMSVD	0.00	0.00	0.00	0.00	RMSVD	0.00	0.00	0.00	0.00	
BIAS	0.00		0.00	0.00	BIAS	0.00		0.00	0.00	
SPD	0.00		0.00	0.00	SPD	0.00		0.00	0.00	
NCMV	0	0	0	0	NCMV	0.00	0.00	0.00	0.00	
NC	0	0	0	0	NC	Ō	0	ō	ō	
					INC	0	0	0	0	

#### Table 5-2 As per Table 5-1, but for WV (B08) rawinsonde statistics

### Table 5-3 As per Table 5-1, but for WV (B09) rawinsonde statistics (a)

(a)					(b)				
	Sele	llite in use : cted channel : et Type :		Himawari-8 B09 (6.9 um) cloudy	( )	Selec	llite in use : sted channel : et Type :		Himawari-9 B09 (6.9 um) cloudy
	ALL REGIONS	NH EX-TROP	TROP	SH EX-TROP		ALL REGIONS	NH EX-TROP	TROP	SH EX-TROP
SPD NCMV	21.95 659198	6.58 -0.14	5.25 -0.20 16.65 414600	5.36 6.29 -0.26 27.38 130821 2921		4.67 5.57 -0.10 22.36 501213 66492	0.09 33.98 89292	5.16	5.90 -0.18 27.16 96253
HIGH LEVEL MVD RMSVD BIAS SPD NCMV NC	4.74 5.67 -0.24 21.85 607361 71598	5.56 6.69 -0.29 36.44 88249	4.45 5.25 -0.22 16.78 399721 51899	5.37 6.30 -0.31 27.75 119391 2798	HIGH LEVEL MVD RMSVD BIAS SPD NCMV NC	4.66 5.56 -0.18 22.29 462325 61361	5.35		
MEDIUM LEVEL MVD RMSVD BIAS SPD NCMV NC	4.94 5.90 0.54 23.35 51837 5231	5.03 6.06 0.49 26.75 25528 4090	4.56 5.18 0.69 10.17 14879 1018	5.22 6.18 0.97 19.00 11430 123	MEDIUM LEVEI MVD RMSVD BIAS SPD NCMV NC	4.80	4.90 5.83 0.84 26.54 20441 4015	4.37 4.96 0.78 9.89 10540 990	4.76 5.67 1.05 17.69 7907 126
LOW LEVEL MVD RMSVD BIAS SPD NCMV NC	$\begin{array}{c} 0.00\\ 0.00\\ 0.00\\ 0.00\\ 0.00\\ 0\\ 0\\ 0\\ 0\\ 0\\ 0\\ 0\\ 0\\ 0\\ 0\\ 0\\ 0\\ $	0.00	0.00 0.00 0.00 0.00 0 0	0.00 0.00 0.00 0.00 0.00 0	LOW LEVEL MVD RMSVD BIAS SPD NCMV NC	0.00 0.00 0.00 0.00 0 0	0.00 0.00 0.00 0.00 0 0	0.00 0.00 0.00 0.00 0 0	0.00 0.00 0.00 0.00 0.00 0

## Table 5-4 As per Table 5-1, but for WV (B10) rawinsonde statistics

(a)					(C)				
~ /	Sele	llite in use : cted channel : et Type :		Himawari-8 B10 (7.3 um) cloudy	( )	Sele	llite in use : cted channel : et Type :		Himawari-9 Bl0 (7.3 um) cloudy
	ALL REGIONS	NH EX-TROP		SH EX-TROP		ALL REGIONS	NH EX-TROP		
RMSVD BIAS SPD NCMV	5.64 -0.34 21.29 699461	5.34 6.46 -0.35 32.83 134100	4.42 5.22 -0.33 16.07 422484	5.25 6.18 -0.46 26.55 142877	ALL LEVEL MVD RMSVD BIAS SPD NCMV	4.63 5.52 -0.23 21.72 536479	5.15 6.16 -0.13 32.03 105325	4.33 5.13 -0.27 15.90 324708	4.97 5.84 -0.30 26.16 106446
NC	78492	22451	52852	3189	NC HIGH LEVEL	69693	23356	43513	2824
HIGH LEVEL MVD RMSVD BIAS SPD NCMV NC	4.71 5.63 -0.37 21.27 602309 70180	6.64 -0.44 35.81	4.42 5.23 -0.35 16.30 396837 51034	5.29 6.22 -0.49 27.39 118825 2924	MVD RMSVD BIAS SPD NCMV	4.63 5.52 -0.29 21.75 463315 60718	6.34 -0.25 35.12 68001	-0.30 16.16 305808	5.89 -0.38 27.21 89506
	4.77 5.74 -0.05 21.43 97152	5.97 -0.10 25.06 47453	4.88 0.12 9.59	4.82 5.69 -0.10 17.25 24052 265	MEDIUM LEVE MVD RMSVD BIAS SPD NCMV NC	4.66 5.53 0.20 21.50 73164 8975	5.71	4.78 0.33 9.61 18900	5.30 0.42 16.54 16940
LOW LEVEL MVD RMSVD BIAS SPD NCMV NC	0.00 0.00 0.00 0.00 0	0.00	0.00 0.00 0.00 0.00 0 0	0.00 0.00 0.00 0.00 0	LOW LEVEL MVD RMSVD BIAS SPD NCMV NC	0.00	0.00	0.00	0.00 0.00 0.00

(a)					(U)					
	Sele	llite in use : cted channel : et Type :		Himawari-8 B13 (10.4 um) cloudy	( )	Sele	Satellite in use : Selected channel : Target Type :			
	ALL REGIONS	NH EX-TROP	TROP	SH EX-TROP		ALL REGIONS				
RMSVD BIAS SPD NCMV NC HIGH LEVEL	5.42 -0.50 18.74 896671 65879 4.72 5.64	5.89 -0.55 24.92 241477 25896 5.47 6.61	5.04 -0.48 14.12 409601 36343	3640 5.52 6.45	ALL LEVEL MVD BIAS SPD NCMV NC HIGH LEVEL MVD RMSVD BIAS	4.40 5.27 -0.45 19.11 698182	4.66 5.63 -0.46 24.62 189953 27533 5.25 6.27 -0.48	$\begin{array}{r} 4.17\\ 4.93\\ -0.45\\ 13.94\\ 323496\\ 30287\\ 4.28\\ 5.06\\ -0.54\\ \end{array}$	4.35 5.18 -0.25 20.54 184733 3183 5.21 6.02 -0.67	
	20.74 383625	34.76 55874	14.88 246792	28.53 80959 2094	SPD NCMV NC	21.25 309796 40078	48005	14.74 196965 26077	28.33 64826 1783	
RMSVD BIAS SPD	4.49 5.41 -0.35 18.92 139376	5.71 -0.37 22.58 61996	3.89 4.48 -0.31 8.77 39416 2288	4.33 5.03 -0.21 15.21 37964 357	RMSVD BIAS SPD	4.40 5.25 -0.31 18.88 106660 11169	5.51 -0.34 21.83 48194	4.32 -0.30 8.58 30475	4.36	
LOW LEVEL MVD RMSVD BIAS SPD NCMV NC	4.42 -0.10 10.29 373670	4.64 -0.48 10.74 123607	4.02 0.78 9.61	3.30 3.89 0.32 9.11 126670 1189	BIAS SPD	4.23 -0.27 10.57 281726	-0.59	3.83 0.56	3.57 0.36	

## Table 5-5 As per Table 5-1, but for IR (B13) rawinsonde statistics

#### 5.2 CSR

The first guess (FG) departure statistics of Himawari-8 and -9 CSR products were evaluated for B08 (6.2  $\mu$ m), B09 (6.9  $\mu$ m) and B10 (7.3  $\mu$ m) in the Himawari-9 IOT phase. The FG value of brightness temperature comes from JMA's Global Spectral Model. Figures 5-1 to 5-3 show FG departure histograms for B08, B9 and B10, respectively. Table 5-6 summarizes FG statistics. Although these statistics represent only single-scene comparisons between Himawari-8 and -9, the values for Himawari-9 (except the mean for B09) are consistent with those of Himawari-8. The Himawari-9 mean for B09 is approximately 0.2 K greater.

Table 5-6 FG departure statistics for Himawari-8 and -9 CSR products (all regions, 02 UTC, 25 February 2017)

Band	Satellite	Ν	Mean [K]	Sdev [K]
band-08 (6.2um)	Himawari-8	53079	1.40	2.04
	Himawari-9	52940	1.42	2.04
band-09 (6.9um)	Himawari-8	50074	1.07	2.06
	Himawari-9	49795	1.29	2.06
band-10 (7.3um)	Himawari-8	42950	0.15	1.70
	Himawari-9	42465	0.17	1.71

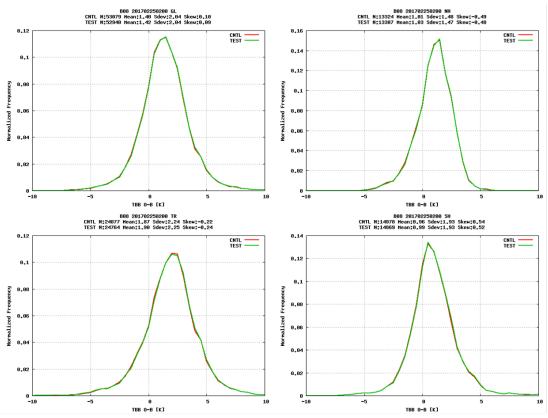


Figure 5-1 FG departure histograms for the B08 (6.2 μm) CSR product in all regions (a), the northern region (b), the tropics (c) and the southern region (d) at 02 UTC on 25 February 2017. CNTL (Himawari-8) and TEST (Himawari-9) are indicated by red and green lines, respectively.

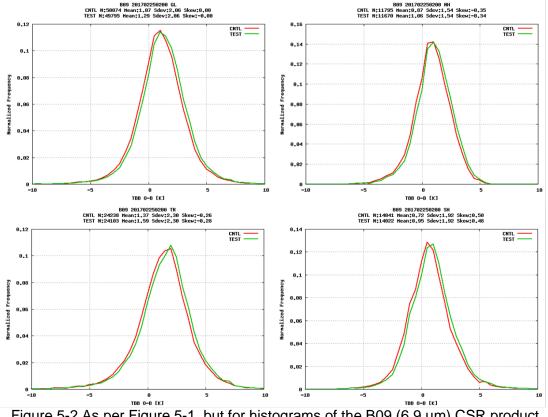


Figure 5-2 As per Figure 5-1, but for histograms of the B09 (6.9  $\mu$ m) CSR product Page 11 of 13

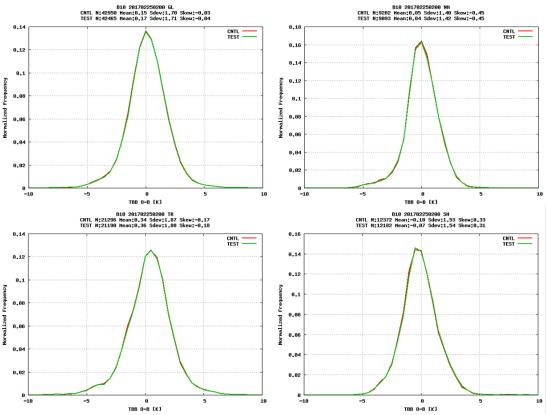


Figure 5-3 As per Figure 5-1, but for histograms of the B10 (7.3 µm) CSR product

#### 6 CONCLUSION

Himawari-9, which features the new-generation AHI visible and infrared imager, was launched on 2 November 2016 and put into in-orbit standby as backup for Himawari-8 on 10 March 2017. This paper reports on the current status of AHI-9's Level-1 and -2 products.

INR errors as determined from observation during the IOT phase are in the same order as those for AHI-8. To validate calibration performance, both GEO-LEO and GEO-GEO approaches were adopted for all 16 bands of AHI-8 and -9. The validation results for AHI-9 calibration generally show close correspondence with those for AHI-8. However, larger biases are seen in several AHI-9 bands (e.g., B01, B06 and B16). One possible reason for the discrepancy in visible and near-infrared bands is that pre-launch calibration slopes were used for AHI-9, whereas coefficients derived from solar diffuser observation are used for AHI-8 bands. In addition to the implementation of new calibration and validation approaches such as a lunar method, calibration trends will also be evaluated further in the near future.

Quality assessment of Himawari-8 and -9 AMVs and CSRs was also carried out. The rawinsonde statistics for AMVs were similar for both satellites. In regard to CSR, the FG departure statistics for B08 and B10 were similar for both satellites, although the B09 FG departure for Himawari-9 was 0.2 K greater.

#### 7 REFERENCES

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